# Monte Carlo Simulations:



**M** onte C arlo S imulations of **T** riple A xis **S** pectrometers

n

V irtual **I**nstrumentation T ool for the E uropean **S** pallation **S** ource

> Klaus Lieutenant Peter Willendrup

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#### **Numerical experiments before Los Alamos**



- In 1777 Georges-Louis Leclerc, comte de Buffon was the first to "numerical experiments" for solving a problem of geometrical probability.
- The experiment involves dropping a needle on a lined surface and can be used to estimate  $\boldsymbol{\pi}$
- In the 1930's, Fermi used sampling methods to estimate quantities involved in controlled fission



 During WW2, "numerical experiments" were applied at Los Alamos for solving mathematical complications of computing fission, criticality, neutronics, hydrodynamics, thermonuclear detonation etc.



- Notable fathers: John v. Neumann Stanislav Ulam Nicholas Metropolis
- Named "Monte Carlo" after Ulam's fathers frequent visits to the Monte Carlo casino in Las Vegas
- Initially "implemented" by letting large numbers of women use tabularized random numbers and hand calculators for individual particle calculations
- Later, analogue and digital computing devices were used





• FERMIAC

ENIAC



Red: 7846 Blue: 2154 Pi: 3.1384



- Los Alamos has since then developed and perfected many different monte carlo codes leading to what is today known as the codes MCNP5 and MCNPX
- State of the art is MCNPX (or soon the merged MCNP6 code) that features numerous (even exotic) particles
- MCNP was originally Monte Carlo Neutron Photon, later N-Particle
- Mainly used for high-energy particle descriptions in weapons, power reactors and routinely used for estimating dose rates and needed shielding
- Does not to date handle coherent scattering of neutrons due to the focus on high energies





#### **Ray-tracing methods**



- When neutrons move in "free space", we use ray-tracing but in most cases in direction source -> detector
- Of course parabolas rather than straight lines are uses to implement gravity



# **Elements of Monte-Carlo raytracing**

- Instrument Monte Carlo methods implement coherent scattering effects
- Uses deterministic propagation where this can be done
- Uses Monte Carlo sampling of "complicated" distributions and stochastic processes and multiple outcomes with known probabilities are involved
- - I.e. inside scattering matter
- Uses the particle-wave duality of the neutron to switch back and forward between deterministic ray tracing and Monte Carlo approach



• Result: A realistic and efficient transport of neutrons in the thermal and cold range



# Simulation codes for Monte Carlo ray-tracing of neutron instruments (incomplete?)

- US codes:
  - NISP (P Seeger & L Daemen, LANL)
  - IDEAS (X-L Wang ORNL, originally H Lee ORNL)
  - Instrument Builder (JK Zhao ORNL)
  - McVine (engine of the DANSE vnf project) (J Lin Caltech)
- European codes:
  - RESTRAX/SIMRES (J Šaroun, NPI & J Kulda, ILL)
  - VITESS (K Lieutenant et al. HZB plus S Manoshin JINR)
  - McStas (P Willendrup & E Knudsen DTU, K Lefmann KU, E Farhi ILL, U Filges PSI)
  - [ NADS (P Bentley ESS) ]
- Japanese code:
  - PHITS (MCNP-like) has a few features for instrument beam transport
- There may be others I do not know of...





#### **Example: Instrument simulation of R2D2**





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## The 'Weight' Parameter



 The trajectories are calculated for single neutrons

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- A highly unequal distribution of a parameter is a challenge
   Choosing most events around the peak(s) and only few in the tails shows poor convergence
  - The method of choice is therefore: weighting the events (according to the distribution)
- The weight parameter can also be used to treat events like reflection
- Weighting is a numerical trick, not the treatment of an ensemble average



### **Absolute Flux values**



• I<sub>CW</sub>= ∫j(λ) d λ

$$\approx (\lambda_{max} - \lambda_{min})/N \sum j(\lambda_i)$$
i
• Iss  $\approx (\lambda_{max} - \lambda_{min})(t_{max} - t_{min})/N \sum j(\lambda_i, t_i)$ 

 By proper normalization, the weight can be expressed as a neutron current (or intensity) [n/ s]

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- Summing of the weights of all trajectories gives the neutron current at any point of the instrument
- By processes like reflection the intensity decreases, while the number of trajectories remains unchanged
- The trajectories are the events of the statistical experiment



#### **VITESS GUI (after loading an instrument)**



# Xconi	rol K	:/				×
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Instru	nen	t R2D2			VITESS 2.9 Click parameter names for hel	p!
Check		input file			Browsel Browsell	
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Stop		random seed 1	_	ge	erator ran3 min. neutron 1.0e-25 gravity off Exit	
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#### **McStas GUI (after loading an instrument)**



McStas: X:/ILL/ILL-Comp/Diffraction/D2B/Diffr1_ILL_part1.instr	
Elle Simulation Neutron site	Help (McDoc)
Instrument file: X:/ILL/ILL-Comp/Diffraction/D2B/Diffr1_ILL_part1.instr	Edit/New Run
Simulation results: <none></none>	Read Plot
Status: Ok	
<pre>HcStas version 1.9.1 - Mar. 29, 2006 Copyright (C) Risce National Laboratory, 1997-2006 Additions (C) Institut Laue Langevin, 2003-2006 All rights reserved Hatlab plotter is NOT available Perl/PGPLOT plotter is NOT available Instrument file: X:/ILL/D11/D11_S4/D11guides.instr Instrument file: X:/ILL/ILL-Comp/Diffraction/D2B/Diffr1_II</pre>	<pre>DEFINE INSTRUMENT My_Instrument(DIST=10) /* Here comes the TRACE section, where the actual */ /* instrument is defined as a sequence of components. */ TRACE /* The Arm() class component defines reference points and orientations */ /* in 3D space. */ COMPONENT Origin = Arm() AT (0,0,0) ABSOLUTE COMPONENT Source = Source_simple(    radius = 0.1, dist = 10, xw = 0.1, yh = 0.1, E0 = 5, dE = 1) AT (0, 0, 0) RELATIVE Origin COMPONENT Emon = E_monitor(    filename = "Emon.dat", xmin = -0.1, xmax = 0.1, ymin = -0.1,</pre>
	<pre>ymax = 0.1, Emin = 0, Emax = 10) AT (0, 0, DIST) RELATIVE Origin COMPONENT PSD = PSD monitor( nx = 128, ny = I28, filename = "PSD.dat", xmin = -0.1, xmax = 0.1, ymin = -0.1, ymax = 0.1) AT (0, 0, 1e-10) RELATIVE Emon /* The END token marks the instrument definition end */ END</pre>







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view from moderator towards sample



# **Parameters in VITESS and McStas**



#### **McStas**

- ID
- criterion 'ray tracing'
- 'colour'
- time of flight [ms]
- weight [n/s]
- location of neutron x [cm]
- location of neutron y [cm]
- location of neutron y [cm]
- wavelength [Å]
- flight direction v<sub>x</sub>/|v|
- flight direction v<sub>y</sub>/<u>|v|</u>
- flight direction vz/|v|
- spin S<sub>x</sub>
- spin S<sub>y</sub>
- spin S<sub>z</sub>

- time of flight [s]
  - weight [n/s]
- location of neutron x [m]
- location of neutron y [m]
- location of neutron y [m]
  - speed v<sub>x</sub>[m/s]
  - speed v<sub>y</sub>[m/s]
  - speed v<sub>z</sub>[m/s]
    - spin S<sub>x</sub>
    - spin S<sub>y</sub>
    - spin S<sub>z</sub>



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#### **Simulation Output**





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#### Example: Al<sub>2</sub>O<sub>3</sub>on PUS







#### **Example: Improvement of R2D2 using NAC sample**





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McStas WITESS





#### Instrument optimization

- New ideas can be tested first in a simulation
- Different option can be compared
- Numerical optimization is possible

#### Virtual experiments

 If some information about the sample is available, it can be checked what time and which settings are needed for the real experiment

#### Instrument and data analysis

- Simulations provide a large amount of information on the properties of the neutrons (e.g. on correlations in phase space and spin space)
- They allow comparing true and measured sample properties

Teaching

E. Farhi, M. Johnson, V. Hugouvieux and W. Kob, ILL Annual Report (2006) 87.







Acknowledgement



# Thank you for your attention

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